

4.4 Hydrogenic atoms

Bohr model^a

Quantisation condition	$\mu r_n^2 \Omega = n\hbar$	(4.71)	r_n n th orbit radius Ω orbital angular speed n principal quantum number (> 0)
Bohr radius	$a_0 = \frac{\epsilon_0 \hbar^2}{\pi m_e e^2} = \frac{\alpha}{4\pi R_\infty} \simeq 52.9 \text{ pm}$	(4.72)	a_0 Bohr radius μ reduced mass ($\simeq m_e$) $-e$ electronic charge Z atomic number h Planck constant $\hbar$ $h/(2\pi)$
Orbit radius	$r_n = \frac{n^2}{Z} a_0 \frac{m_e}{\mu}$	(4.73)	E_n total energy of n th orbit ϵ_0 permittivity of free space m_e electron mass
Total energy	$E_n = -\frac{\mu e^4 Z^2}{8\epsilon_0^2 \hbar^2 n^2} = -R_\infty \hbar c \frac{\mu}{m_e} \frac{Z^2}{n^2}$	(4.74)	α fine structure constant μ_0 permeability of free space
Fine structure constant	$\alpha = \frac{\mu_0 c e^2}{2\hbar} = \frac{e^2}{4\pi\epsilon_0 \hbar c} \simeq \frac{1}{137}$	(4.75)	
Hartree energy	$E_H = \frac{\hbar^2}{m_e a_0^2} \simeq 4.36 \times 10^{-18} \text{ J}$	(4.76)	E_H Hartree energy
Rydberg constant	$R_\infty = \frac{m_e c \alpha^2}{2\hbar} = \frac{m_e e^4}{8\hbar^3 \epsilon_0^2 c} = \frac{E_H}{2\hbar c}$	(4.77)	R_∞ Rydberg constant c speed of light
Rydberg's formula ^b	$\frac{1}{\lambda_{mn}} = R_\infty \frac{\mu}{m_e} Z^2 \left(\frac{1}{n^2} - \frac{1}{m^2} \right)$	(4.78)	λ_{mn} photon wavelength m integer $> n$

^aBecause the Bohr model is strictly a two-body problem, the equations use reduced mass, $\mu = m_e m_{\text{nuc}} / (m_e + m_{\text{nuc}}) \simeq m_e$, where m_{nuc} is the nuclear mass, throughout. The orbit radius is therefore the electron–nucleus distance.

^bWavelength of the spectral line corresponding to electron transitions between orbits m and n .



Hydrogenlike atoms – Schrödinger solution^a

Schrödinger equation		
$-\frac{\hbar^2}{2\mu}\nabla^2\Psi_{nlm}-\frac{Ze^2}{4\pi\epsilon_0 r}\Psi_{nlm}=E_n\Psi_{nlm} \quad \text{with} \quad \mu=\frac{m_em_{\text{nuc}}}{m_e+m_{\text{nuc}}} \quad (4.79)$		
Eigenfunctions		
$\Psi_{nlm}(r,\theta,\phi)=\left[\frac{(n-l-1)!}{2n(n+l)!}\right]^{1/2}\left(\frac{2}{an}\right)^{3/2}x^le^{-x/2}L_{n-l-1}^{2l+1}(x)Y_l^m(\theta,\phi) \quad (4.80)$		
with $a=\frac{m_e}{\mu}\frac{a_0}{Z}$, $x=\frac{2r}{an}$, and $L_{n-l-1}^{2l+1}(x)=\sum_{k=0}^{n-l-1}\frac{(l+n)!(-x)^k}{(2l+1+k)!(n-l-1-k)!k!}$		
Total energy	$E_n=-\frac{\mu e^4 Z^2}{8\epsilon_0^2 \hbar^2 n^2} \quad (4.81)$	E_n total energy ϵ_0 permittivity of free space
Radial expectation values	$\langle r \rangle = \frac{a}{2}[3n^2 - l(l+1)] \quad (4.82)$	h Planck constant m_e mass of electron
	$\langle r^2 \rangle = \frac{a^2 n^2}{2}[5n^2 + 1 - 3l(l+1)] \quad (4.83)$	$\hbar$ $h/2\pi$ μ reduced mass ($\simeq m_e$)
	$\langle 1/r \rangle = \frac{1}{an^2} \quad (4.84)$	m_{nuc} mass of nucleus Ψ_{nlm} eigenfunctions
	$\langle 1/r^2 \rangle = \frac{2}{(2l+1)n^3 a^2} \quad (4.85)$	Ze charge of nucleus $-e$ electronic charge
Allowed quantum numbers and selection rules ^b	$n=1,2,3,\dots \quad (4.86)$	L_p^q associated Laguerre polynomials ^c
	$l=0,1,2,\dots,(n-1) \quad (4.87)$	a classical orbit radius, $n=1$
	$m=0,\pm 1,\pm 2,\dots,\pm l \quad (4.88)$	r electron–nucleus separation
	$\Delta n \neq 0 \quad (4.89)$	Y_l^m spherical harmonics
	$\Delta l = \pm 1 \quad (4.90)$	a_0 Bohr radius $=\frac{\epsilon_0 \hbar^2}{\pi m_e e^2}$
$\Delta m = 0 \quad \text{or} \quad \pm 1 \quad (4.91)$		
$\Psi_{100}=\frac{a^{-3/2}}{\pi^{1/2}}e^{-r/a} \quad \Psi_{200}=\frac{a^{-3/2}}{4(2\pi)^{1/2}}\left(2-\frac{r}{a}\right)e^{-r/2a}$		
$\Psi_{210}=\frac{a^{-3/2}}{4(2\pi)^{1/2}}\frac{r}{a}e^{-r/2a}\cos\theta \quad \Psi_{21\pm 1}=\mp\frac{a^{-3/2}}{8\pi^{1/2}}\frac{r}{a}e^{-r/2a}\sin\theta e^{\pm i\phi}$		
$\Psi_{300}=\frac{a^{-3/2}}{81(3\pi)^{1/2}}\left(27-18\frac{r}{a}+2\frac{r^2}{a^2}\right)e^{-r/3a} \quad \Psi_{310}=\frac{2^{1/2}a^{-3/2}}{81\pi^{1/2}}\left(6-\frac{r}{a}\right)\frac{r}{a}e^{-r/3a}\cos\theta$		
$\Psi_{31\pm 1}=\mp\frac{a^{-3/2}}{81\pi^{1/2}}\left(6-\frac{r}{a}\right)\frac{r}{a}e^{-r/3a}\sin\theta e^{\pm i\phi} \quad \Psi_{320}=\frac{a^{-3/2}}{81(6\pi)^{1/2}}\frac{r^2}{a^2}e^{-r/3a}(3\cos^2\theta-1)$		
$\Psi_{32\pm 1}=\mp\frac{a^{-3/2}}{81\pi^{1/2}}\frac{r^2}{a^2}e^{-r/3a}\sin\theta\cos\theta e^{\pm i\phi} \quad \Psi_{32\pm 2}=\frac{a^{-3/2}}{162\pi^{1/2}}\frac{r^2}{a^2}e^{-r/3a}\sin^2\theta e^{\pm 2i\phi}$		

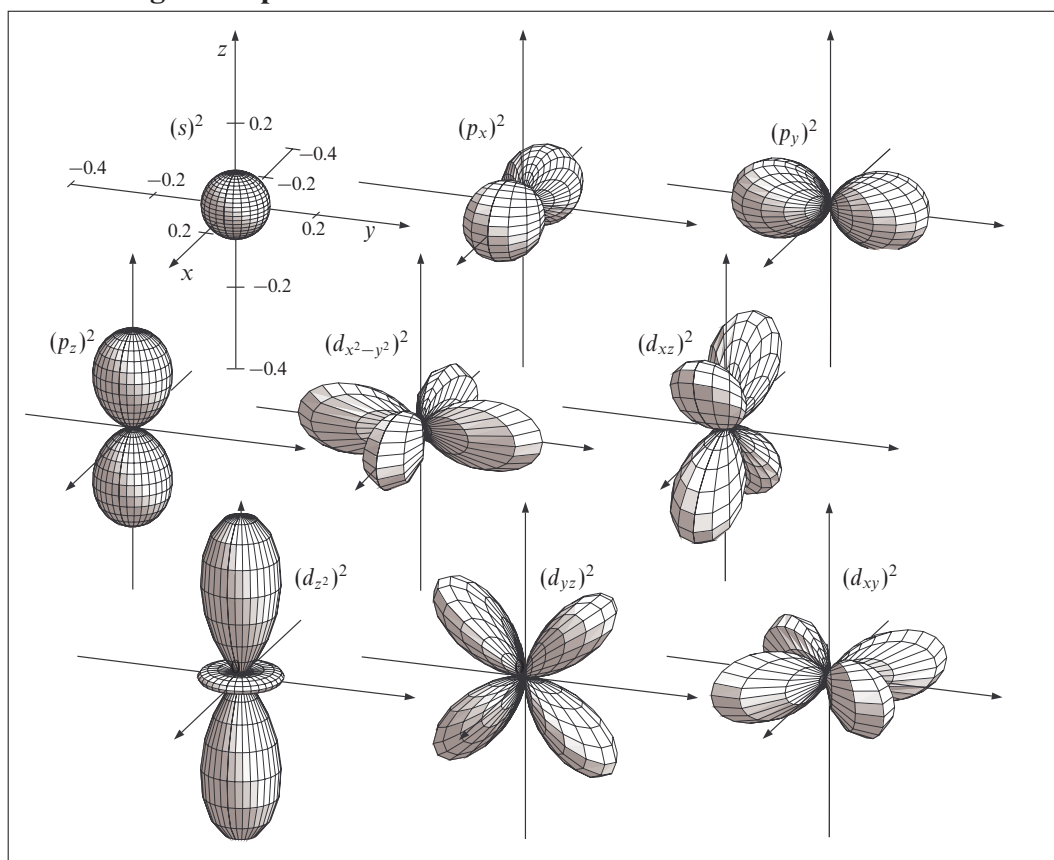
^aFor a single bound electron in a perfect nuclear Coulomb potential (nonrelativistic and spin-free).

^bFor dipole transitions between orbitals.

^cThe sign and indexing definitions for this function vary. This form is appropriate to Equation (4.80).



Orbital angular dependence



s orbital ($l=0$)	$s = Y_0^0 = \text{constant}$	(4.92)	Y_l^m spherical harmonics ^a θ, ϕ spherical polar coordinates
p orbitals ($l=1$)	$p_x = \frac{-1}{2^{1/2}}(Y_1^1 - Y_1^{-1}) \propto \cos \phi \sin \theta$	(4.93)	
	$p_y = \frac{i}{2^{1/2}}(Y_1^1 + Y_1^{-1}) \propto \sin \phi \sin \theta$	(4.94)	
	$p_z = Y_1^0 \propto \cos \theta$	(4.95)	
d orbitals ($l=2$)	$d_{x^2-y^2} = \frac{1}{2^{1/2}}(Y_2^2 + Y_2^{-2}) \propto \sin^2 \theta \cos 2\phi$	(4.96)	
	$d_{xz} = \frac{-1}{2^{1/2}}(Y_2^1 - Y_2^{-1}) \propto \sin \theta \cos \theta \cos \phi$	(4.97)	
	$d_{z^2} = Y_2^0 \propto (3 \cos^2 \theta - 1)$	(4.98)	
	$d_{yz} = \frac{i}{2^{1/2}}(Y_2^1 + Y_2^{-1}) \propto \sin \theta \cos \theta \sin \phi$	(4.99)	
	$d_{xy} = \frac{-i}{2^{1/2}}(Y_2^2 - Y_2^{-2}) \propto \sin^2 \theta \sin 2\phi$	(4.100)	

^aSee page 49 for the definition of spherical harmonics.

